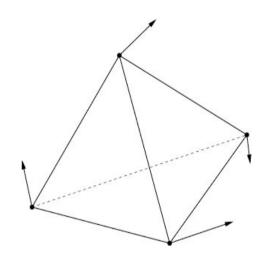
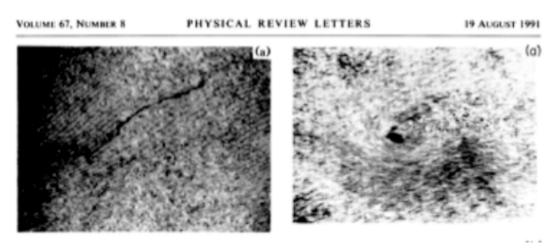
# **Geometry and statistics in turbulence**



Alain Pumir, CNRS and Université de Nice.

Fluid turbulence occurs when fluid motion is fast (large Reynolds). It is characterized by :

- A complex ('turbulent') spatio temporal dynamics.
- The existence of a wide range of spatially excited scales (... notion of scaling).
- Despite its complexity, turbulent flows exhibit welldefined structures. Examples of vortex tubes :

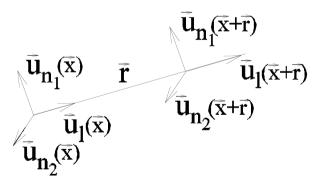


Douady et al, PRL 1991

Roughly speaking, most investigations of turbulence consider separately either of the two aspects:



dependence as a function of r of the structure functions:



the local structure

(i.e., the
geometry) of intermittent
regions in the flow (such as
vorticity filaments).

→ Aim here : capture both aspects, by investigating multipoint correlation functions.

#### **Objective of the work:**

Develop a theoretical understanding and a description of the fluctuating velocity field that captures both the scaling and structural aspects of the flow.

#### An important remark:

To properly characterize the flow, focus on the full velocity gradient tensor:

$$m_{ab} = \partial_a u_b$$

Or its coarse grained generalization:

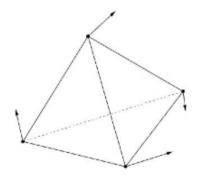
$$M_{ab} = \frac{1}{V_R} \times \int_{V_R} m_{ab} d^d r$$

The lagrangian approach has turned out to be extremely useful, in particular for the solution of the Kraichnan model (see, e.g.: Shraiman and Siggia, Nature 2000, and Falkovich et al. Rev. Mod. Phys., 2001).

#### It pays to just follow the flow!!

At the minimum, 4 points are needed to construct any finite difference approximation of the velocity derivative tensor, ( $\sim$  M).

=> A **tethrahedron** is the minimal structure one has to study.



The evolution of the tetrahedron and M can be modelled by a stochastic differential equation (Chertkov et al, 1999) which we are studying directly.

#### **Potential pay-offs:**

- ✓ Fundamental information about the nonlinear processes in the Navier-Stokes equations.
- ✓ Invitation to think about multipoint correlation.
- ✓ Get insight about the transfer process between scales (Pumir et al, 2001, Bandi et al, 2006).
- ✓ Potentially, particle based LES (Shraiman et al, 2003).

~ new way to think about the turbulence problem (a related approach : Chevillard + Meneveau, 2006, 2007)

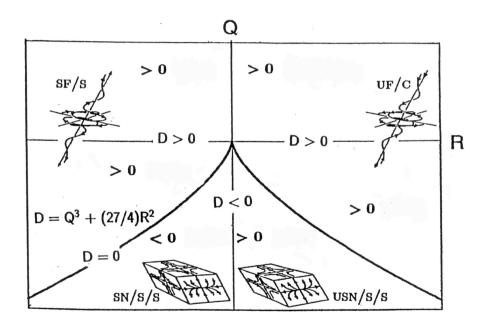
# M as a diagnostic of flow topology

- \* The eigenvalues of M characterize the local topology of the flow.
- \* They depend (Cayley-Hamilton) on the two invariants:

$$Q = -\frac{1}{2}Tr(m^2)$$

$$R = -\frac{1}{3}Tr(m^3)$$

#### Vortex dominated



Strain dominated

#### Outline of the presentation

- The stochastic M-model: derivation and definition.
- Semi-classical solutions of the model.
- Numerical solutions and comparisons with DNS with an isotropic forcing.
- Numerical solutions in the presence of a large shear flow.
- Recent experimental results and new questions.
- Conclusions and perspectives.

#### The stochastic model:

Derivation and definition

#### The stochastic M-model: derivation and definition (1)

Write the Navier-Stokes equation for the velocity gradient tensor :

$$\frac{dm_{ab}}{dt} + m_{ab}^2 = -\partial_{ab}p + viscosity + forcing$$

Crucial ingredient : the pressure hessian

 Isotropic approximation (restricted Euler dynamics, cf Vieillefosse, Cantwell):

$$\partial_{ab} p = -\frac{1}{3} Tr(m^2) \delta_{ab}$$

The resulting system can be completely solved, with the help of the invariants Q and R (Q =  $-tr(m^2)/2$ ; R=  $-tr(m^3)/3$ ):

-> Finite time singularity!

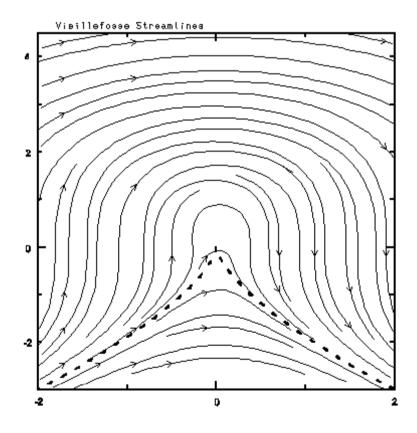
## Singularity of the restricted Euler model

Streamlines of the vector field :

$$dR/dt=Q^2/2$$
  
 $dQ/dt=-R/3$ 

Divergence along the separatrix line

$$4Q^3 + 27 R^2 = 0$$



#### The stochastic M-model: derivation and definition (2)

 To go beyond the Vieillefosse singularity, one needs to introduce the geometry of the Lagrangian set of points.

Equation for the geometry, derived from:

$$\frac{d\rho}{dt} = v = \rho . M + \xi$$

Where :  $\rho M$  = coherent component of the velocity field (k~1/R)  $\xi$  = rapidly fluctuation component (k >> 1/R).

 $\rho_i^a = (\vec{\rho}_i)^a$  = set of reduced coordinates, parametrizing the tetrad.

Introduce the **moment of inertia tensor** :  $g = \rho^t \rho$ 

#### The stochastic M-model: derivation and definition (3)

 Equation for the coarse-grained velocity gradient tensor (obtained from an approximation of the pressure Hessian, based on analytical and numerical results):

$$\frac{dM}{dt} + (M^2 - \Pi Tr(M^2)) = \alpha (M^2 - \Pi Tr(M^2)) + \eta$$
 
$$(M^2 - \Pi Tr(M^2)) \qquad \text{``elocal ``omponent of the pressure}$$
 
$$\alpha (M^2 - \Pi Tr(M^2)) \qquad \text{``elocal ``omponent of the pressure.}$$
 
$$\eta \qquad \text{``fluctuating component}$$
 
$$\left(\Pi = \frac{g^{-1}}{Tr(g^{-1})}\right)$$

Reduction of the nonlinearity through the pressure Hessian : the importance of this effect is measured by

#### The stochastic M-model: derivation and definition (4)

One finally obtains the following system of stochastic differential equations:

$$\begin{cases} \frac{dM}{dt} + (1 - \alpha)(M^2 - \Pi Tr(M^2)) = \eta \\ \frac{d\rho}{dt} - g.M - M^t.g - \beta \sqrt{Tr(MM^t)}(g - Tr(g)Id) = 0 \end{cases}$$

The effect of the noise in the g-equation is assumed to (mostly) restore the isotropy of the g-tensor. It is substituted here by the  $\beta$ -term.

The noise  $\eta$  is modelled by a Gaussian white noise term, obeying the K41-scaling ( $\rho^2 = \text{Tr}(g)$ ):

$$\langle \eta_{ab}(\rho, t).\eta_{cd}(0, 0) \rangle = \gamma \left( \delta_{ac} \delta_{bd} - \frac{1}{3} \delta_{ab} \delta_{cd} \right) \frac{\varepsilon}{\rho^2} \delta(t)$$

#### The stochastic M-model: derivation and definition (5)

 Summary: the model thus reduces to a set of nonlinear, stochastic differential equations, with 3 dimensionless parameters:

- Reduction of nonlinearity by the parameter C.
- Strength of the isotropy restoring term (for the g tensor),
   B.
- Intensity of the fluctuations in the M-equation, γ.

#### **Energy balance**

Define the energy at scale  $\rho$  by E=Tr(VV<sup>t</sup>)/2 by :  $V_i^a = \rho_i^a M_{ha}$ 

#### Equation of evolution of the energy:

$$\partial_{t}E(\rho) = -\frac{\partial}{\partial \rho_{i}^{a}} \left\langle V_{i}^{a} tr(VV^{T}) \right\rangle_{\rho} + \alpha \left\langle tr(VV^{T}M) \right\rangle_{\rho} + (coupling \ with \ small \ scales)$$

• Physical interpretation : 
$$-\frac{\partial}{\partial \rho_i^a} \left\langle V_i^a tr(VV^T) \right\rangle_{\hat{\rho}} \text{ large scale energy flux}$$

$$\alpha \langle tr(VV^TM) \rangle_{\rho}$$
 : eddy-damping term

(see Borue and Orszag, 1998, Meneveau and Katz 2000,...)

The model provides a way to compute the statistical properties of the M-tensor as a function of scale!

What is the qualitative behavior of the solutions of this system of equations?

N.b.: it depends on the three parameters:  $\alpha$ ,  $\beta$  and  $\gamma$ .

#### Methods of resolution

of the system

#### The equation satisfied by the Eulerian PDF...

A Fokker-Planck equation for the Eulerian PDF can be derived from this stochastic system:

$$\partial_t P(M,g,t) = L.P(M,g,t)$$

The stationary solutions must satisfy the system:

$$L.P = 0$$

$$\int dMP(M,g) = 1$$

$$\int dM P(M,g) = 1$$
 
$$P(M,g = L^2 Id) \approx \exp\left[-\frac{Tr(MM^T)}{\left(\varepsilon L^{-2}\right)^{2/3}}\right]$$
 (Gaussian distribution at the integral scale)

#### ... and its solution in terms of path integrals

The system can be solved using Green's functions methods:

$$P(M,g) = \int dM' \int dT G_{-T}(M,g \mid M',g') P(M',g')$$

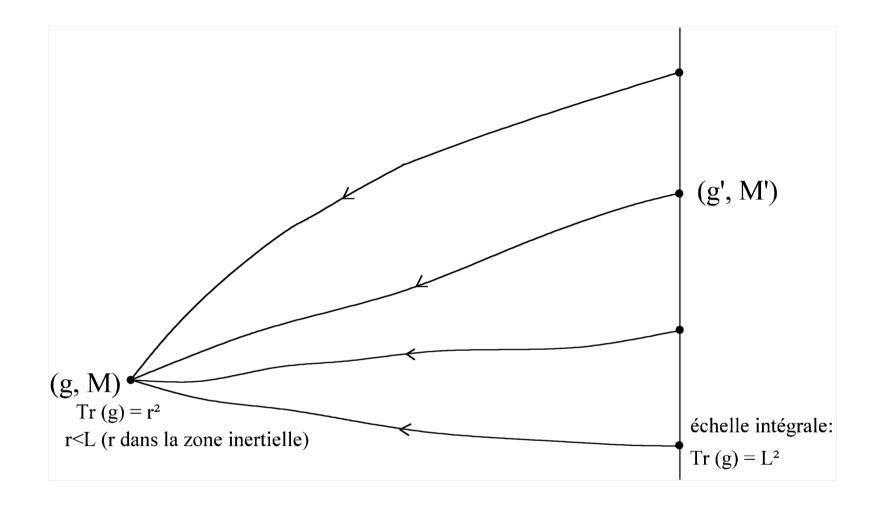
(G : Green's function; P(M',g') : boundary condition)

With:

$$G_{-T}(M,g \mid M',g') = \int [DM''][Dg''] \exp[-S(M'';g'')]$$

Hence:

$$P(M,g) = \int dM' \int dT \int \left[DM''\right] \int \left[Dg''\right] \exp\left[-S(M'';g'') + Tr(M'M'^t)/(\varepsilon L^{-2})^{2/3}\right]$$
(Green's function) (boundary condition)



Starting from an initial condition at the integral scale, one integrates the system up to a fixed scale r (in the inertial range). In principle, one has to *integrate over all trajectories* in phase space.

#### (Approximate) method of resolution (1)

One could use a straightforward **Monte-Carlo method** (exact in principle)

#### Difficulty:

the method is **extremely inefficient**, since one has to deal with trajectories with widely different statistical weight (by orders of magnitude!).

Obtaining reliable numerical results requires prohibitively large computer time.

0<sup>th</sup> order approximation : look for **deterministic solutions** (γ=0)

-> encouraging results when compared with DNS (Chertkov et al, 1999)

#### (Approximate) method of resolution (2)

Use here the **semiclassical approximation** (saddle point approximation of the path integral)

Method: one considers only the trajectory for which the action is minimal (the one with the largest statistical weight).

Hope: The method should provide important information, especially since many trajectories do not contribute very much.

**Drawback**: the method is not rigorous; it is difficult to control the errors made.

=> A better algorithm has to be implemented to understand the effect of fluctuations (~Monte-Carlo), and to really estimate the errors made by using the semi-classical approximation.

# Numerical solutions of the system in the semiclassical approximation with isotropic forcing.

#### Comparison with DNS data

A. Naso and A. Pumir, Phys. Rev. E 72, 056318 (2005)

### Scaling laws of the 2nd and 3rd order moments of M:

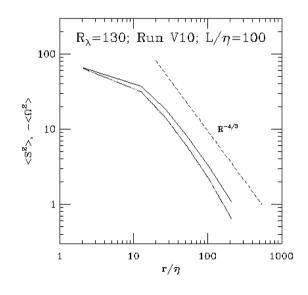
DNS solutions ( $R_{\lambda}$ =130; 256<sup>3</sup>)

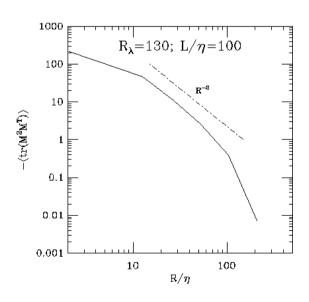
According to the K41 scaling laws,

$$\langle \Delta u(r) \rangle \propto r^{1/3}$$
 so

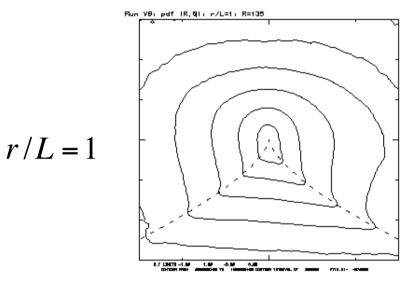
$$\langle M(r) \rangle \propto r^{-2/3}$$
 and  $\langle \omega^2 \rangle, \langle Tr(S^2) \rangle \propto r^{-4/3}$   $\langle -Tr(M^2M^t) \rangle \propto r^{-2}$ 

**DNS results**: these three quantities follow the expected Kolmogorov scaling

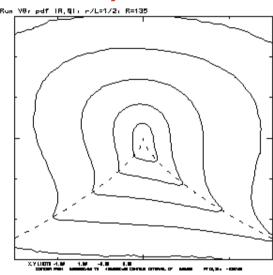




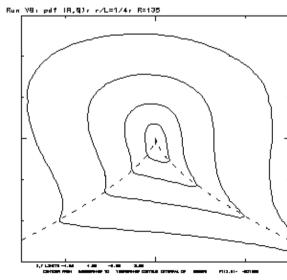
#### Evolution of P(R,Q) as a function of scale; DNS solutions ( $R_{\lambda}$ =130; 256<sup>3</sup>)



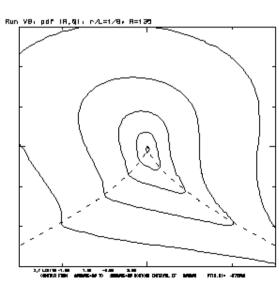
$$r/L = \frac{1}{2}$$



$$r/L = \frac{1}{4}$$



$$r/L = \frac{1}{8}$$

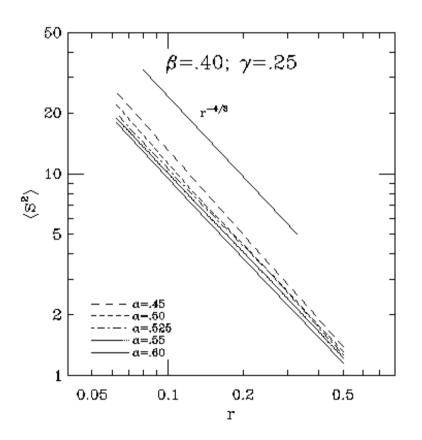


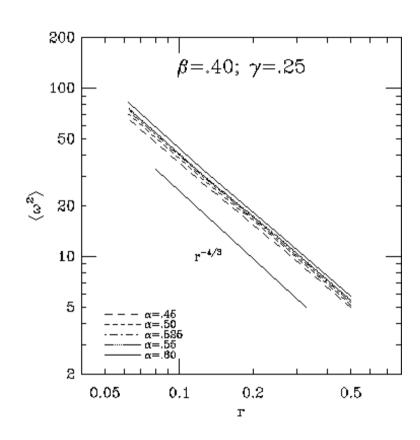
#### Model predictions

The parameter that has the most important effect on the solution is α
 (reduction of the nonlinearity).

• The predictions of the model agree with DNS results provided  $\alpha$  is in a narrow interval around  $\alpha \sim 0.5$ .

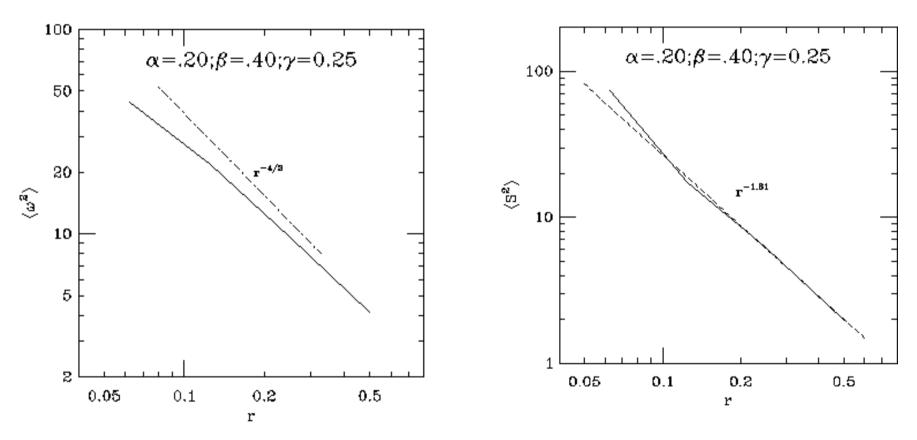
### Scaling properties of the matrix M: model results (1)





The second moment of M has the right scaling provided  $\alpha$  is not too small !

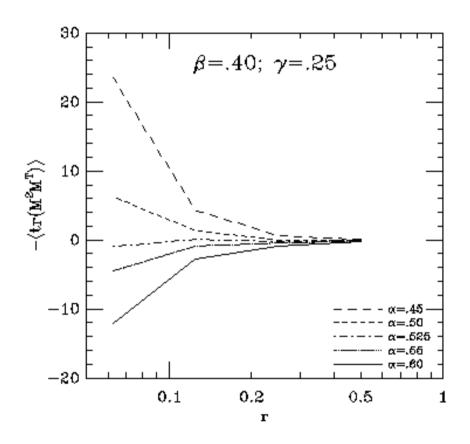
### Scaling properties of the matrix M: model results (2)



At small value of  $\alpha$ , the strain grows with a power that differs significantly from 4/3

The 'reduction of nonlinearity' should not be too small!

### Scaling properties of the matrix M: model results (3)



The sign of  $\langle tr(M^2M^t) \rangle$  is negative, as it should, for small values of  $\alpha$ .

The 'reduction of nonlinearity' should not be too large!

### Scaling properties of the matrix M: model results (4)

Influence of the parameter  $\beta$ :

Not much effect provided  $\beta$  is large enough.

Influence of the parameter  $\gamma$ :

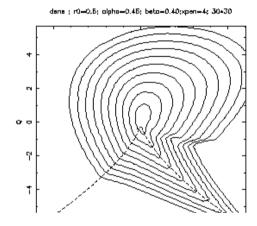
Main effect: change the numerical value of

$$\langle \omega^2 \rangle \times r^{4/3}$$

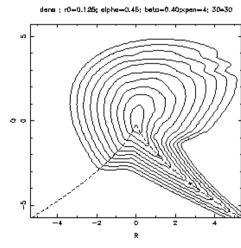
### Evolution of P(R,Q) as a function of scale; semiclassical solutions of the model (1)

Parameters :  $\alpha = 0.45$ ;  $\beta = 0.4$ ;  $\gamma = 0.25$ 

$$\frac{r}{L} = \frac{1}{2}$$

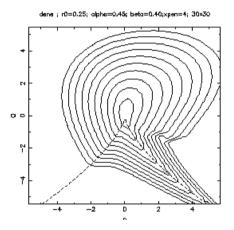


$$\frac{r}{L} = \frac{1}{8}$$

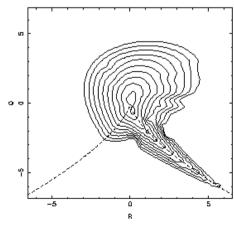


$$\frac{r}{L} = \frac{1}{4}$$

$$\frac{r}{L} = \frac{1}{16}$$



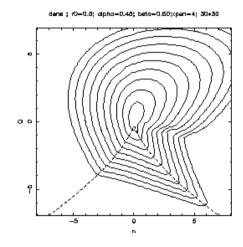
dene ; r0=0.0825; alpha=0.45; beta=0.40; pen=4; pen=4



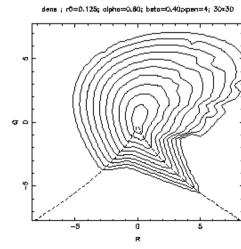
#### **Evolution of P(R,Q) as a function of scale;** semiclassical solutions of the model (2)

Parameters :  $\alpha = 0.6$ ;  $\beta = 0.4$ ;  $\gamma = 0.25$ 

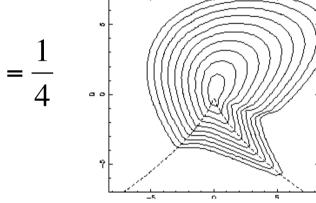
$$\frac{r}{L} = \frac{1}{2}$$



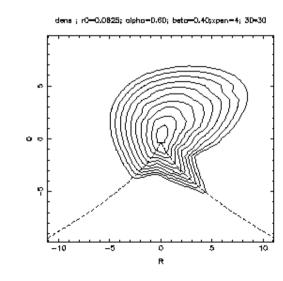
$$\frac{r}{L} = \frac{1}{8}$$



$$\frac{r}{L} = \frac{1}{4}$$



$$\frac{r}{L} = \frac{1}{16}$$

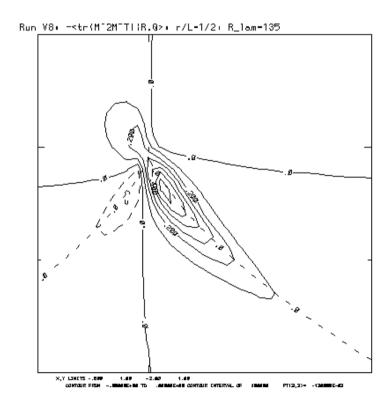


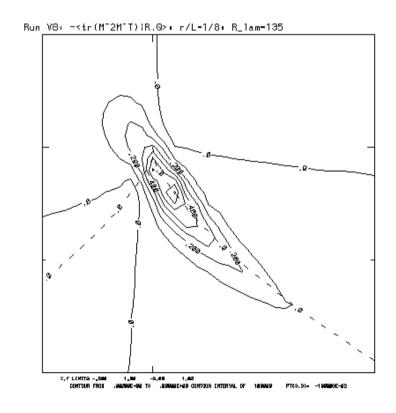
dene : r0-0.25; giphq-0.60; betq-0.40;xpen-4; 30+30

### Scale dependence of the energy transfer density : DNS

$$\frac{r}{L} = \frac{1}{2}$$

$$\frac{r}{L} = \frac{1}{8}$$



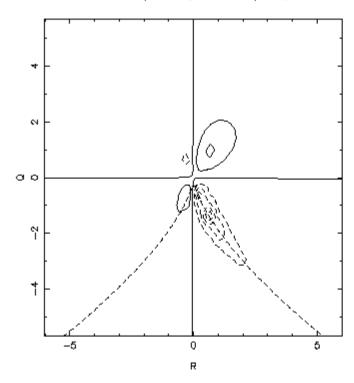


### Scale dependence of the energy transfer density: semiclassical solution

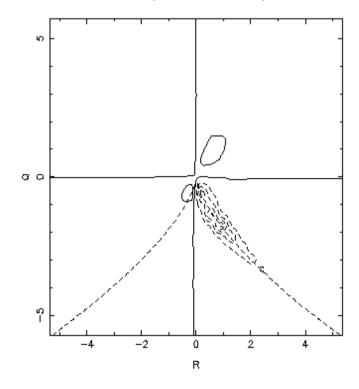
$$\frac{r}{L} = \frac{1}{2}$$

$$\frac{r}{L} = \frac{1}{8}$$

trnsf; r0=0.5; alpha=0.45; beta=0.40;xpen=4; 30\*30



trnsf; r0=0.125; alpha=0.45; beta=0.40;xpen=4; 30\*30



#### Summary : acceptable values of $\alpha$

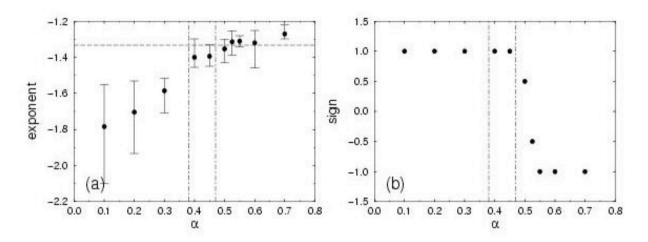
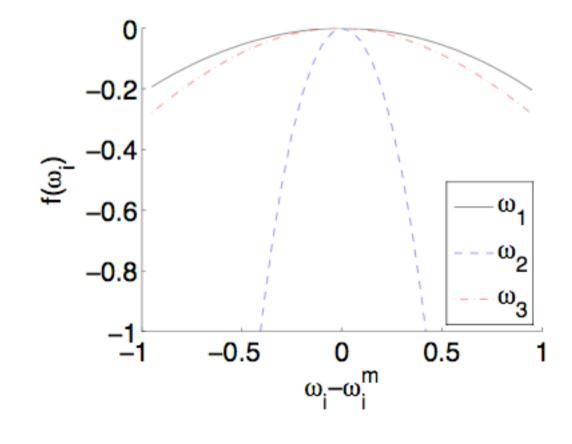


FIG. 5: Dependance with respect to  $\alpha$  of (a) the scaling law exponent of  $\langle S^2 \rangle$  and (b) the sign of the energy transfer  $\langle -r^2Tr(M^2M^t) \rangle$  ( $\beta$ =0.4,  $\gamma$ =0.25). In (a) the dashed line indicates the Kolmogorov prediction -4/3. In (b) is plotted for each value of  $\alpha$  the average of the sign of the energy transfer at the different scales considered. The range of values of  $\alpha$  leading to a qualitatively acceptable behavior of the model's solutions is delimited by the vertical dot-dashed lines.

The solution is acceptable provided a is in a narrow interval around  $\alpha \sim 0.4-0.5$ !

# Semiclassical solution : a caveat

- What we have done:
   determine the
   optimal solution, and
   ignore the
   contributions of other
   nearby trajectories.
- Effect of varying vorticity around the optimum: a better calculation (~MC) is necessary (A. Naso et al, 2007).



# Numerical solutions of the system in the semiclassical approximation with a large scale shear.

A. Naso, M. Chertkov and A. Pumir, J. Turb. (2006)

### The issue of return to isotropy

- One of the postulates of turbulence theory is the universality of small scale velocity fluctuations, which implies that as the scale r diminishes, the flow properties should restore isotropy.
- Study here an homogeneous shear flow.
- Nb: Experimental data (Shen and Warhaft, 2000) and numerical data (Pumir&Shraiman, 1995,1996) suggest that the return to isotropy is much slower than naively expected.

#### The problem studied here

- The tetrad model can be used to study several kinds of forcing, simply by changing the large scale condition.
- -> impose a large scale shear, and calculate the scale dependence of P(R,Q), and other quantities.
- Same equations as in the isotropic case; simply change the large scale boundary condition :

$$P(M.g = L^{2}Id) \sim \exp\left[\frac{-Tr\left[(M-\Sigma)(M-\Sigma)^{t}\right]}{(\varepsilon L^{-2})^{2/3}}\right]$$

Where : 
$$\Sigma = \begin{bmatrix} 0 & s & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}$$
; s measures the **shear intensity**

## Scale dependence of P(R,Q): semiclassical solutions with s=0,1,6

$$\frac{r}{L} = \frac{1}{2}$$

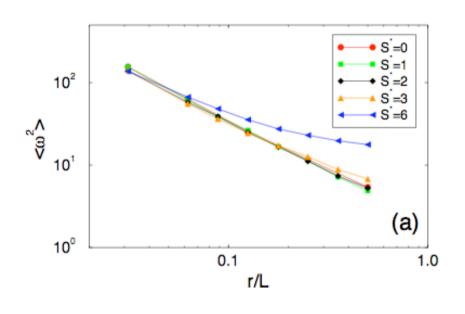
$$\frac{r}{L} = \frac{1}{4}$$

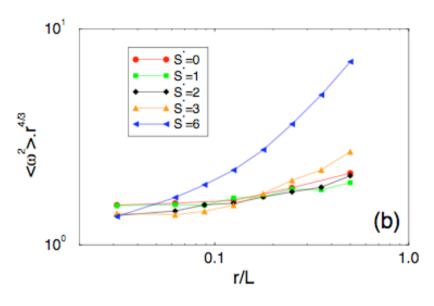
$$\frac{r}{L} = \frac{1}{8}$$

$$\frac{r}{L} = \frac{1}{16}$$

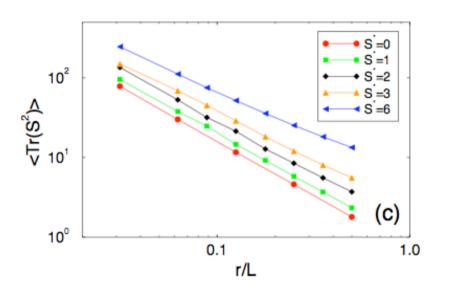
Parameters :  $\alpha$ =0.6,  $\beta$ =0.4;  $\gamma$ =0.25

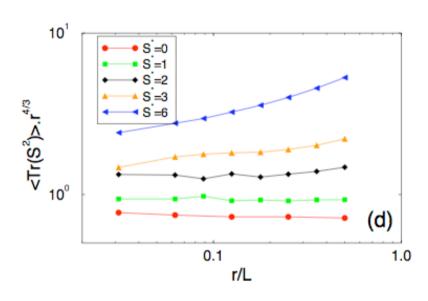
## Scale dependence of $<\omega^2>$ at different values of s



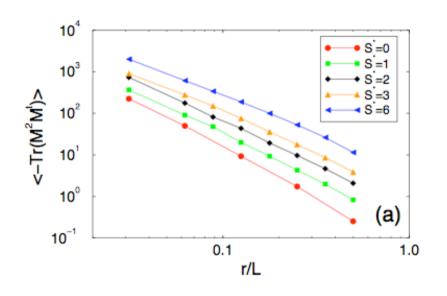


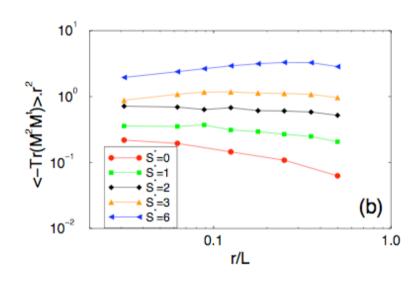
## Scale dependence of <Tr(S)<sup>2</sup>> at different values of s





### Scale dependence of the energy transfer at different values of s





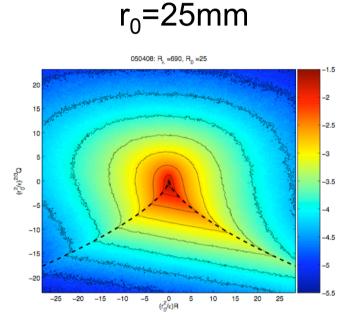
### The issue of return to isotropy

- Our results are consistent with the accepted view that the effects of large scale anisotropy decrease when the scale decreases.
- New finding: difference of behavior between vorticity dominated and strain dominated structures. The anisotropy effects decrease faster for vorticity dominated quantities (enstrophy) rather than for strain dominated objects (strain, energy transfer).
- Faster relaxation of vorticity dominated quantities towards isotropy may be consistent with the facts that vorticity is found to be more intense, hence less sensitive to the large scale forcing.

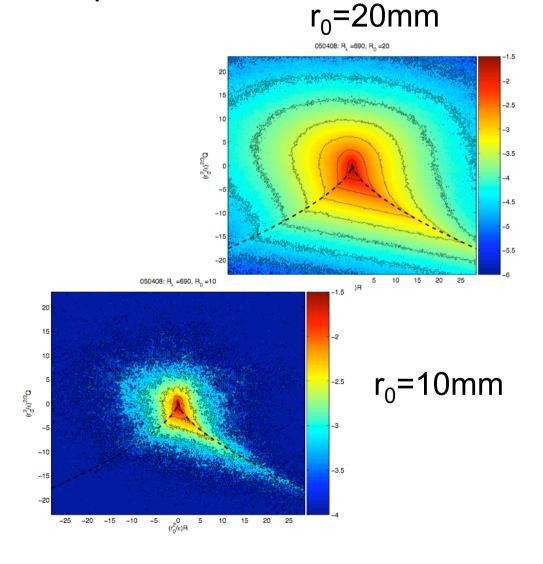
# New experimental developments and new questions.

### Experimental results : pdf(R,Q)

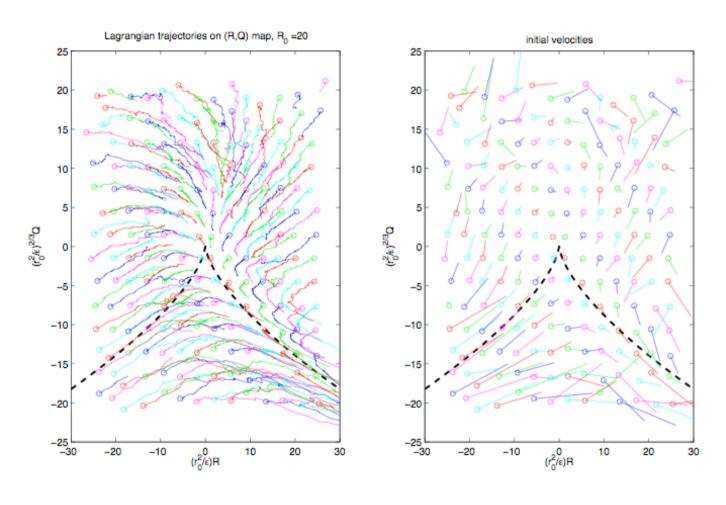
L=70mm;  $\eta$ =0.03mm



H. Xu and E. Bodenschatz Preliminary results



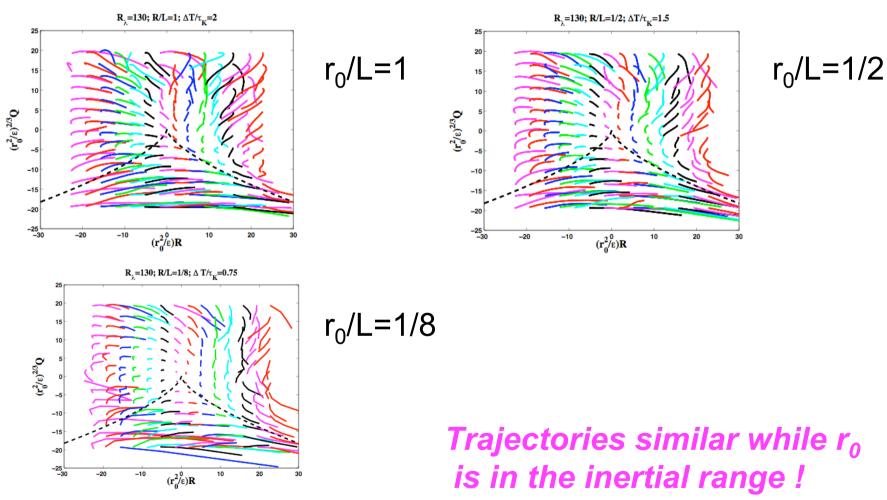
# Experimental results: Trajectories in the (R,Q) plane



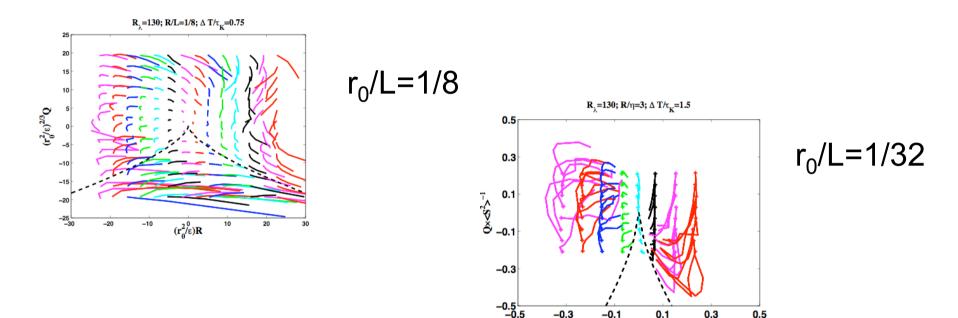
 $R_{\lambda} = 690$ ;  $r_0 = 20$ mm

H. Xu and E. Bodenschatz

### Numerical results: Trajectories in the (R,Q) plane



### Numerical results: Trajectories in the (R,Q) plane

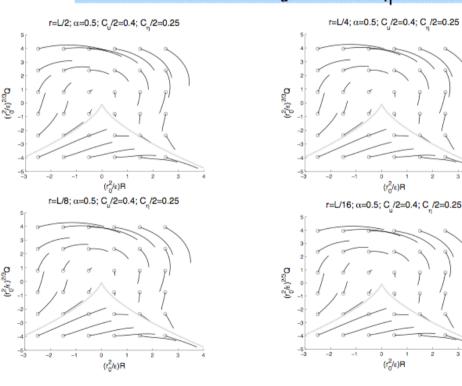


The character of the trajectories changes when r0 in the dissipative range (here,  $\eta/L \sim 100$ )!

n.b.: Difference with Chevillard + Meneveau (2006,2007)

### Model predictions: Trajectories in the (R,Q) plane

r variable;  $\alpha$ =0.5;  $C_u/2$ =0.4;  $C_n/2$ =0.25



- ✓ Little dependence on space  $r_0$
- ✓ Qualitative similarities.
- ✓ Quantitative differences (alignment properties not correctly taken into account)

### Conclusions and perspectives.

### Conclusions and perspectives (1)

- Our work is based on a dynamical model of turbulent velocity fluctuations, that contains several key fluid mechanical ingredients.
- The model is formulated in terms of a stochastic differential equations, that depend on 3 dimensionless parameters.
- The solutions have been obtained in the semiclassical limit, in two cases.
  - <u>isotropic forcing</u>: comparison with DNS results shows the important role of the **nonlinearity reduction** (role of the parameter  $\alpha$ ).
  - <u>anisotropic forcing</u>: difference in the properties of return to isotropy between *vorticity* dominated and *strain* dominated structures.

### Conclusions and perspectives (2)

- Easy to study the influence of boundary conditions at large scales on small scales.
- In progress: development of an hybrid method that incorporates more precisely the fluctuations in the dynamics (... beyond the semiclassical approximation). Expected output: find out about the importance of the fluctuations as a function of the flow structures.
- Very recent development : new experimental results from the Göttingen (also Zürich, Risø and Lyon) group

exciting new developments expected...

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